

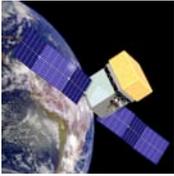
Studies of EGRET unidentified sources with a novel image restoration technique

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GLAST Analysis of Extended Source



- **GLAST Point Spread Function**
 - » **Varies from 0.1° to 10° (2 order of magnitude)**
 - Depends on energy and incident angle.
 - » **Worse than most extended sources**
- **Image restoration is essential for extended source analysis.**

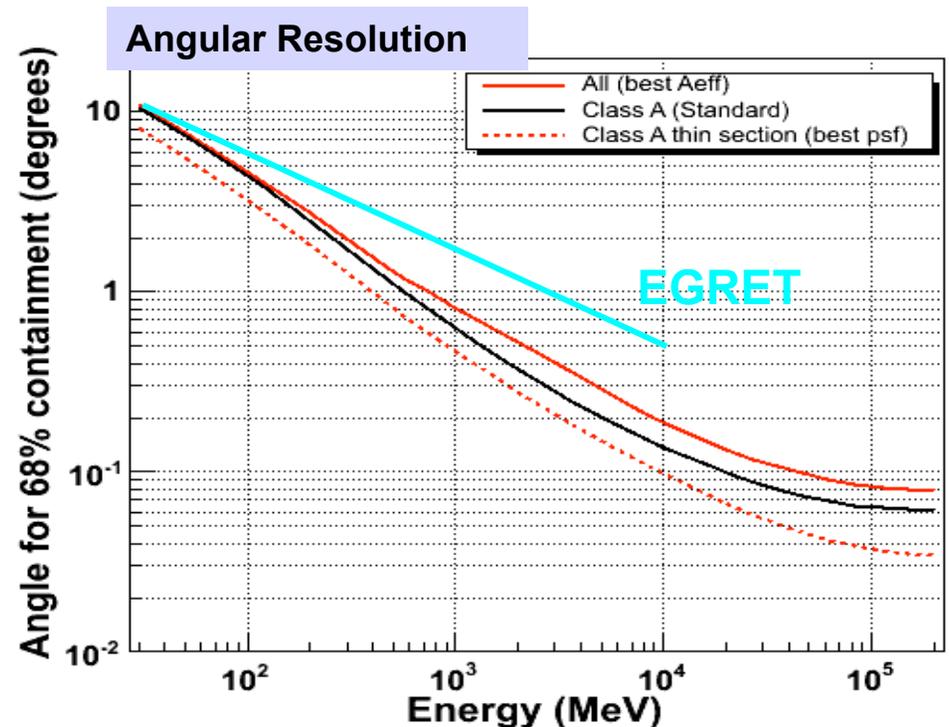




Image Restoration Technique

- **Richardson-Lucy Deconvolution Algorithm.**

- » **When we observe an event at position x .**

- $P(x:\xi)$: probability that it came from a “true” position ξ due to instrument response.

$$\psi^{r+1}(\xi) = \int \tilde{\phi}(x) \frac{\psi^r(\xi) P(x:\xi)}{\int P(x:\zeta) \psi^r(\zeta) d\zeta} dx$$

Lucy 1974

Richardson 1974

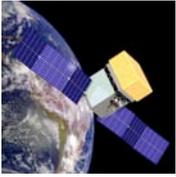
- » **Generalization to Event-by-event $P_k(x:\xi)$.**

$$\psi^{r+1}(\xi) = \frac{1}{N} \psi^r(\xi) \sum_{k=1}^N \frac{P_k(x_k:\xi)}{\int P_k(x_k:\zeta) \psi^r(\zeta) d\zeta}$$

- **Can be used for event-by-event data with varying PSF.**
- No energy spectrum assumption necessary.
- » **Point sources can be incorporated using dual-channel method.**

$$\psi = \psi_{\text{point}} + \psi_{\text{extended}}$$

Hook&Lucy 1994



Wavelet Denoising



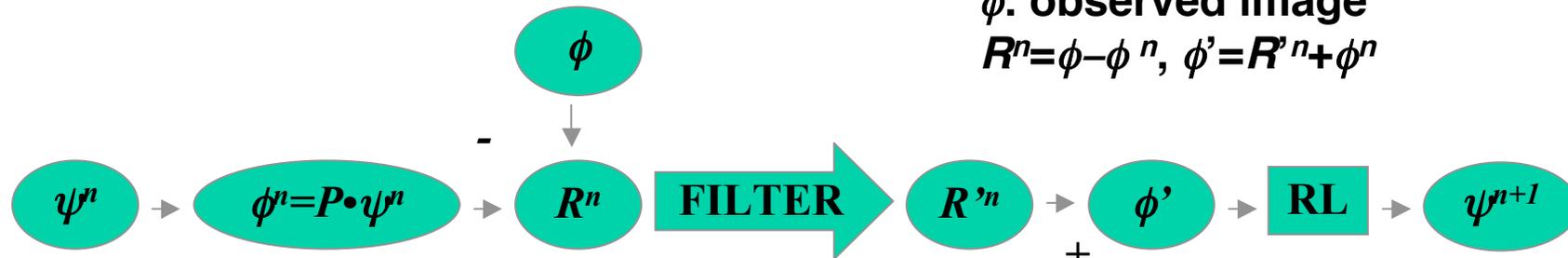
- Minimize the effect of Poisson noise.

Starck&Murtagh 1994

- » Wavelet filtering technique.

ϕ : observed image

$$R^n = \phi - \phi^n, \quad \phi' = R'^n + \phi^n$$



Wavelet decomposition

$$\psi = \sum_{j=0}^n w_j + c_n$$

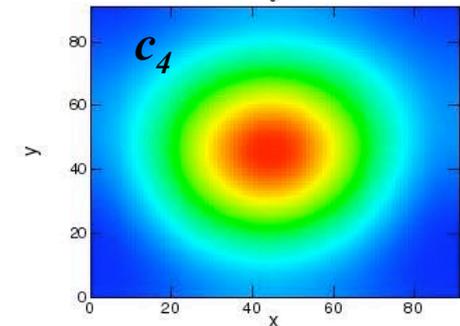
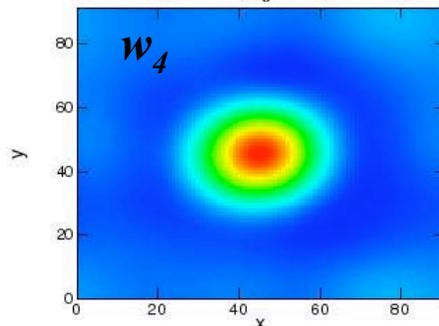
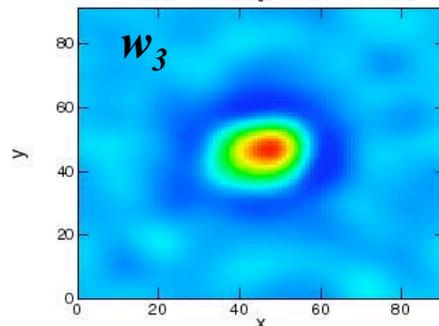
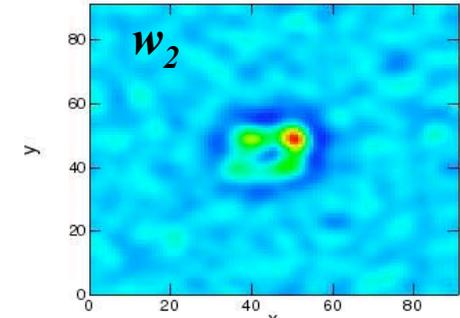
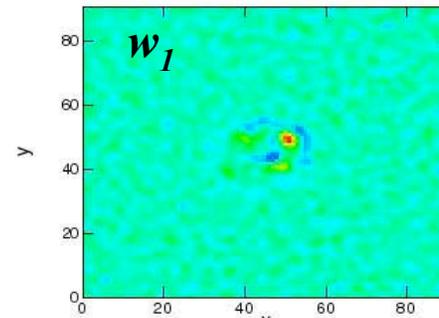
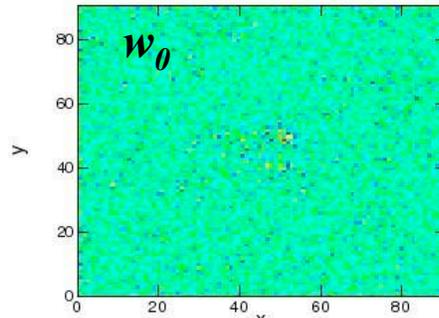


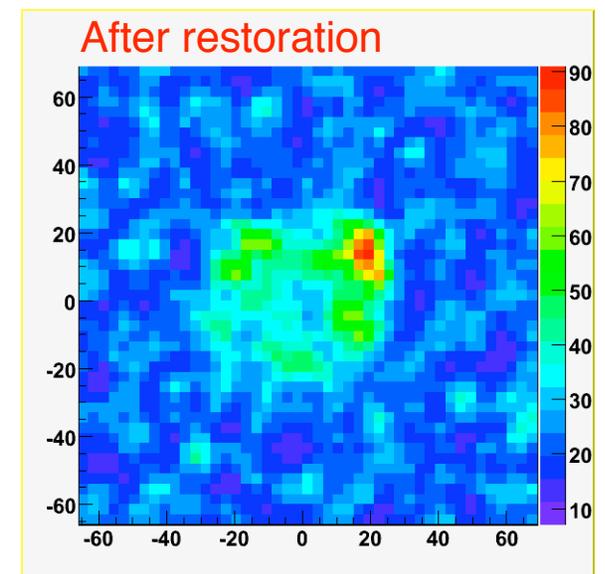
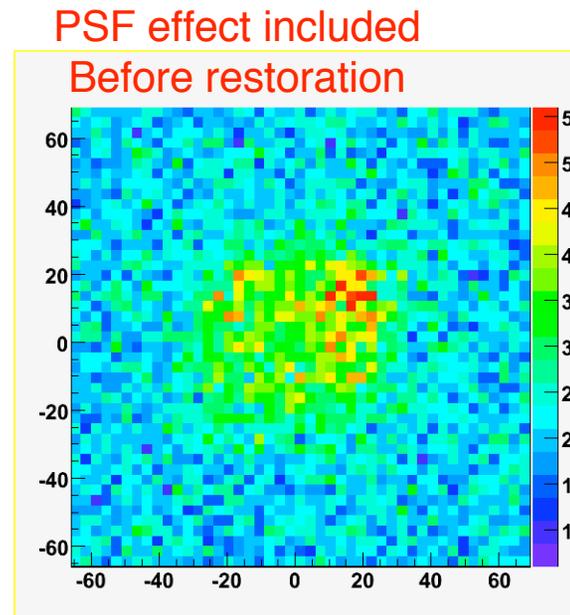
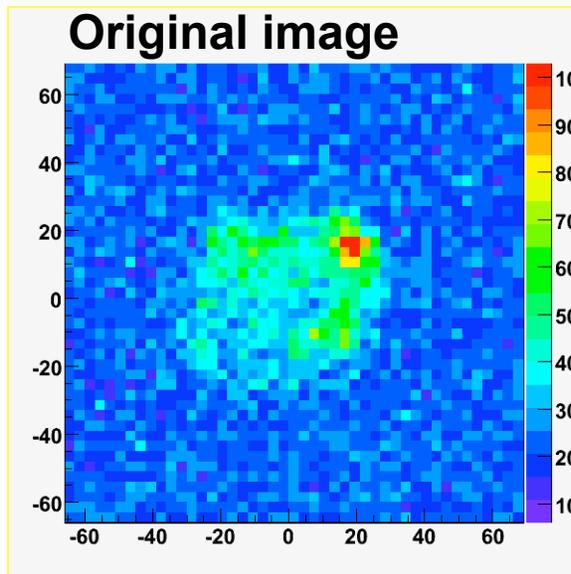
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Image Restoration of Toy MC



- **Restored image gives better representation of input image.**
 - » **NW rim clearly stands out.**
 - » **Overall shape recovered.**
 - » **Poor image at low statistics.**
 - Image restoration can not fix statistical fluctuation

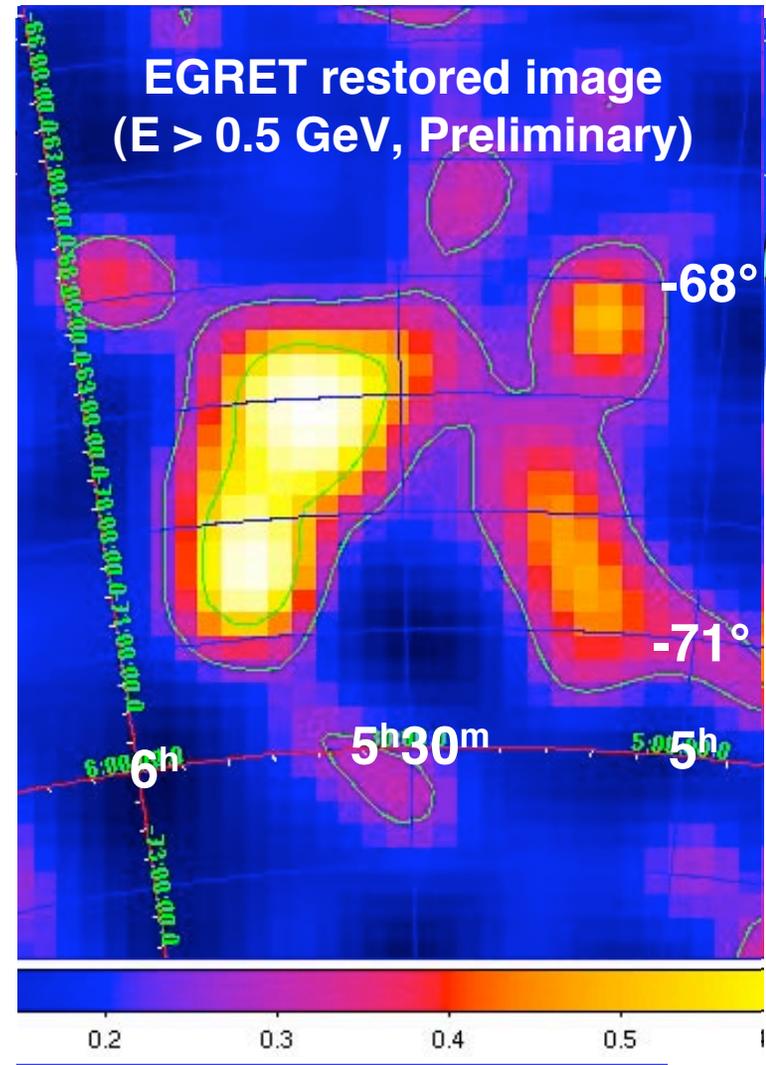
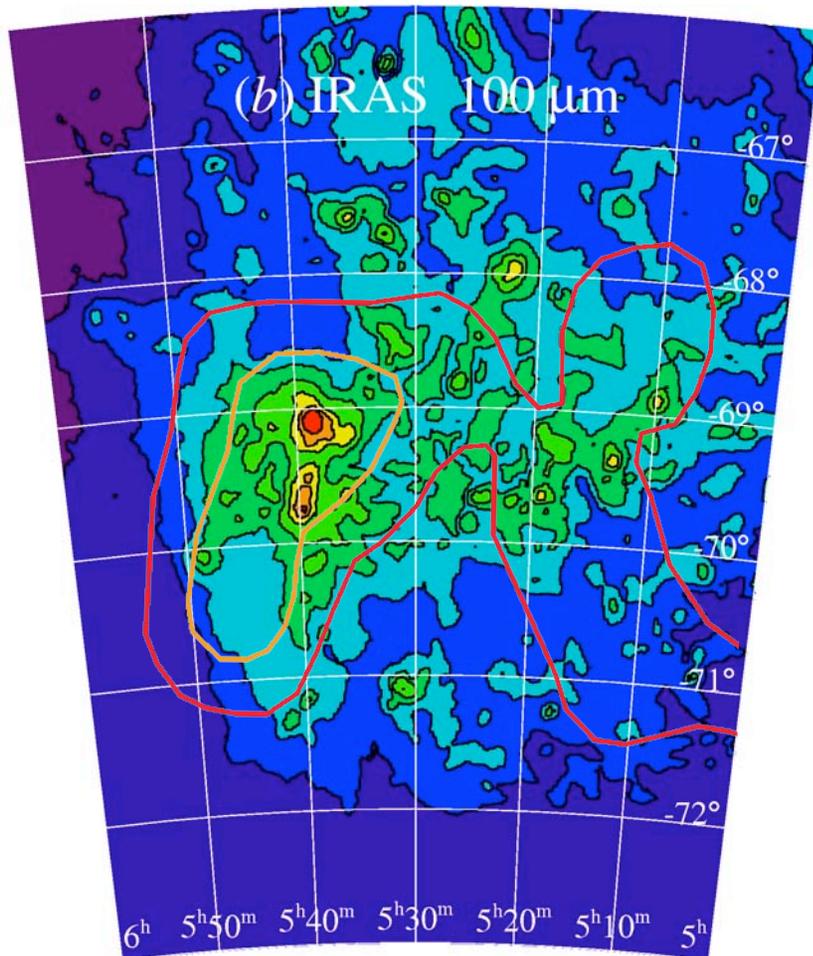




Test Case with known EGRET sources



- **Large Magellanic Cloud**

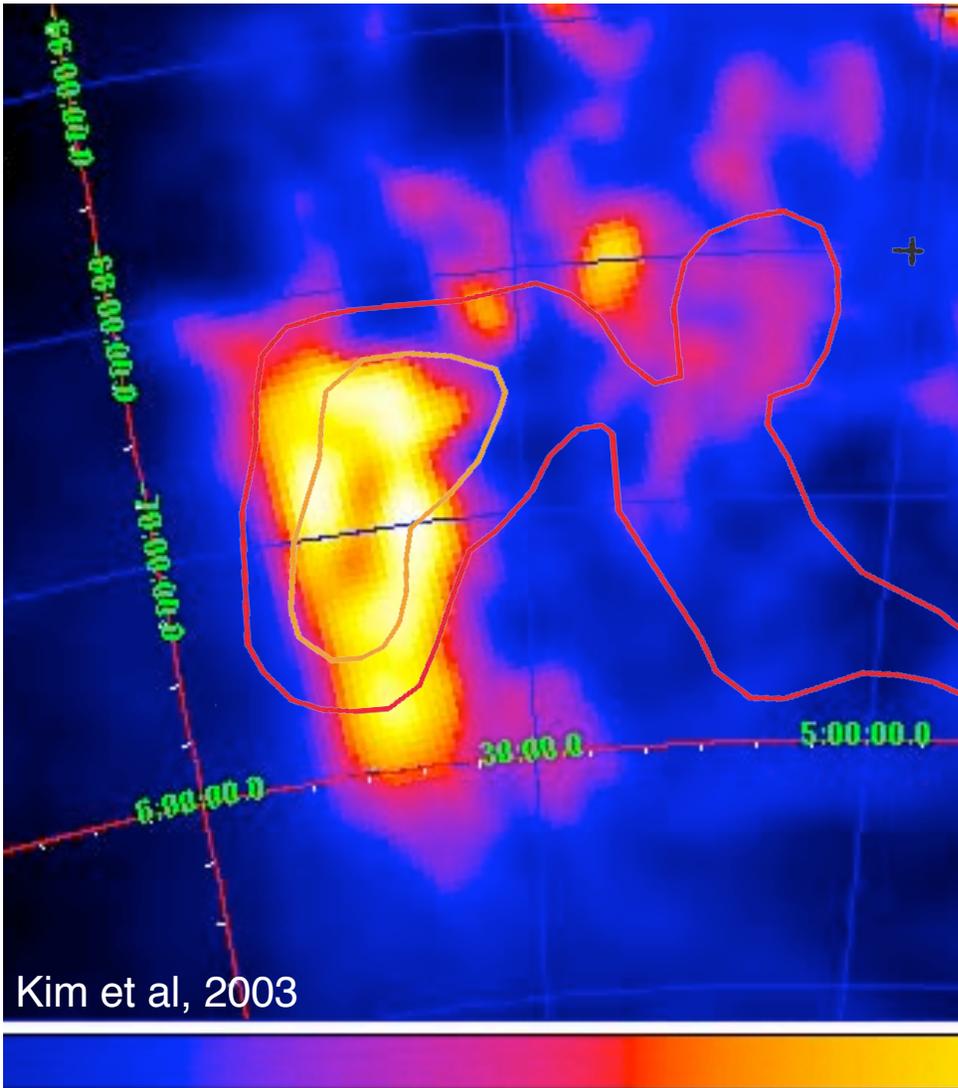


Seth Digel

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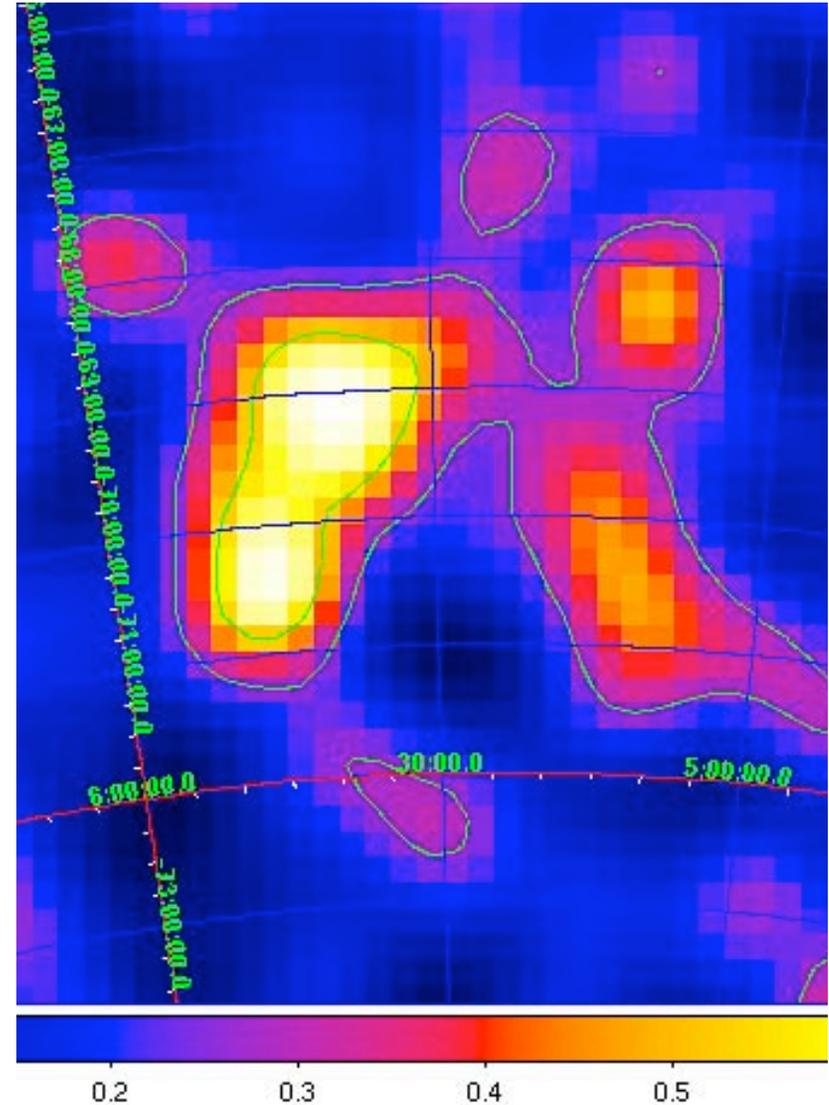


Comparison with HI Map for LMC



Kim et al, 2003

1E+21 2E+21 3E+21



0.2 0.3 0.4 0.5



Kookaburra Region



- Two HESS sources are too close to resolve.

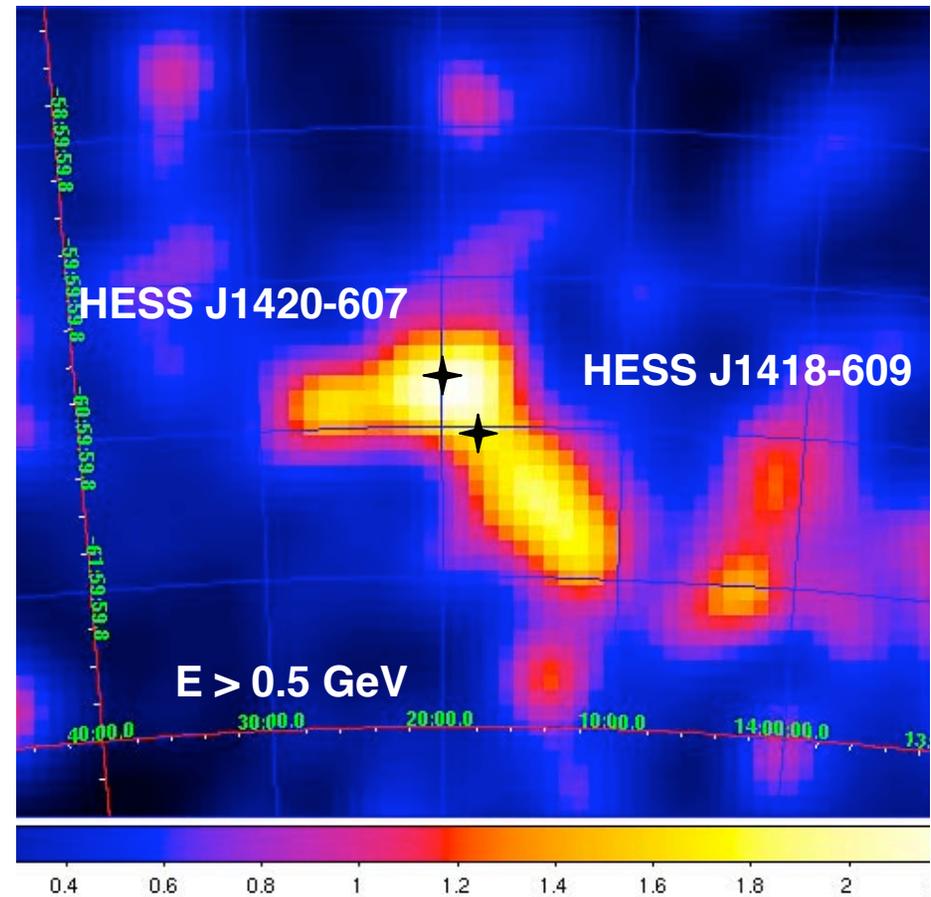
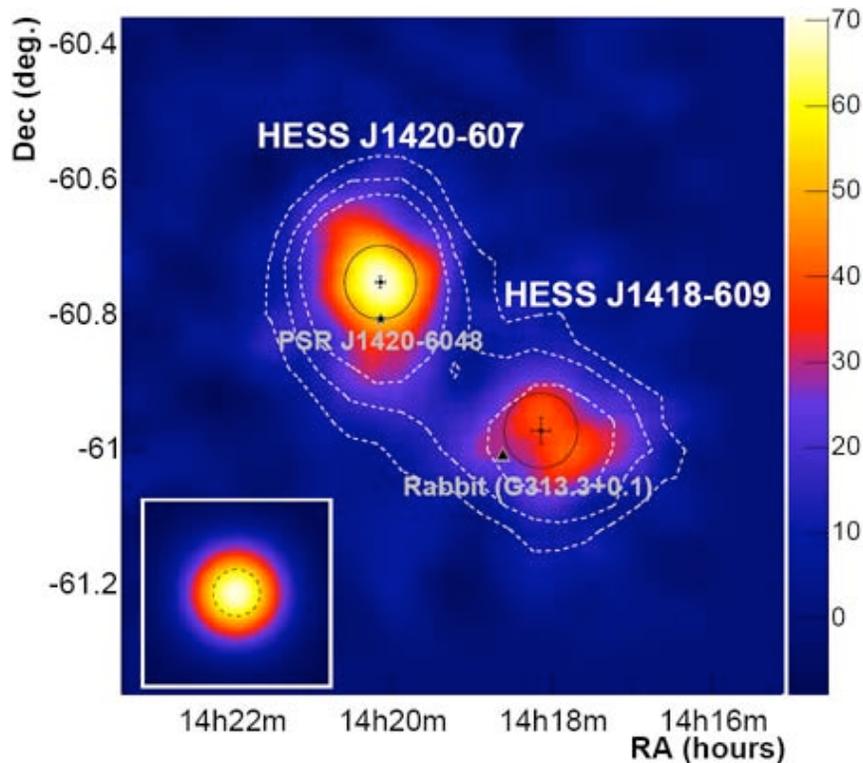


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High Galactic Latitude EGRET UID



- **3EG J1234-1318**
 - » An extended source? Two point sources?
 - » Source position is offset from brightest spot.

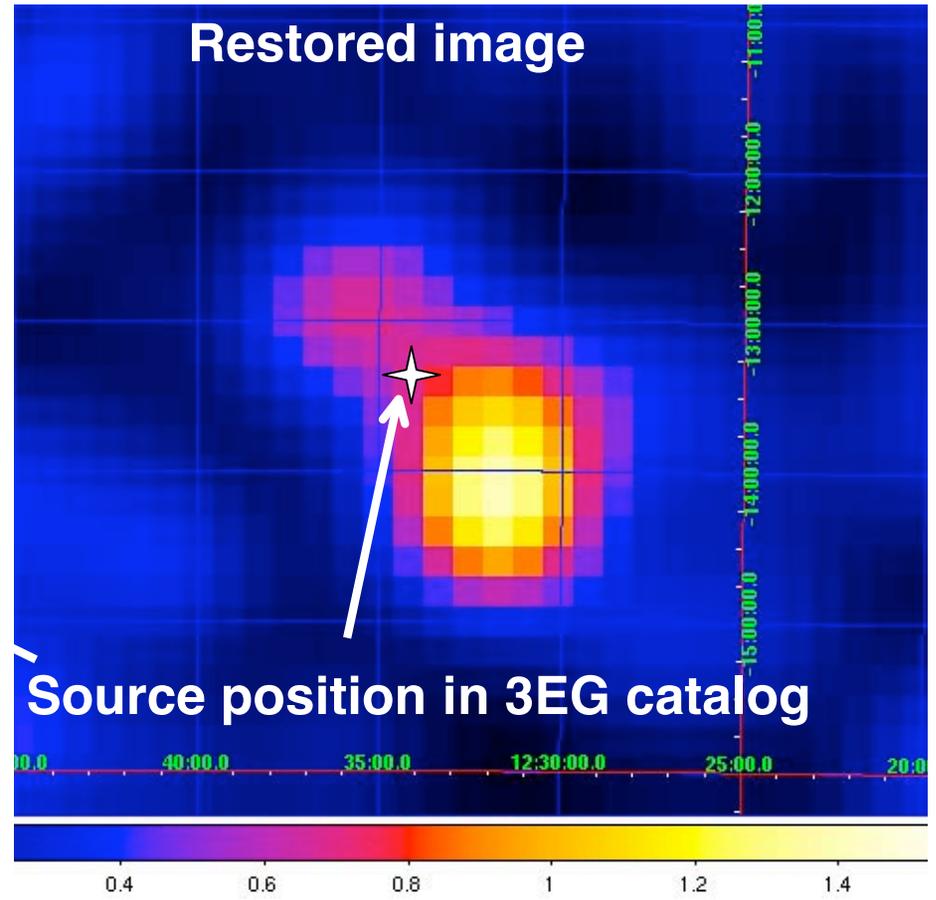
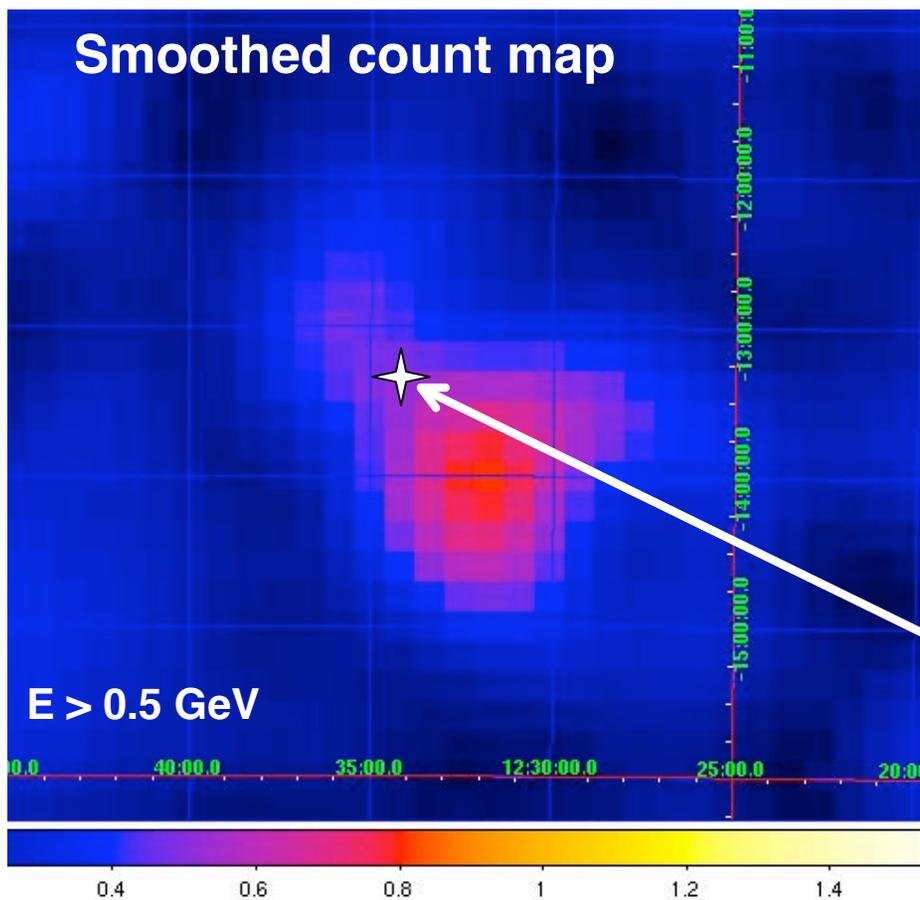


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Galactic Center



- Point source (from 3EG catalog) removal works very well.

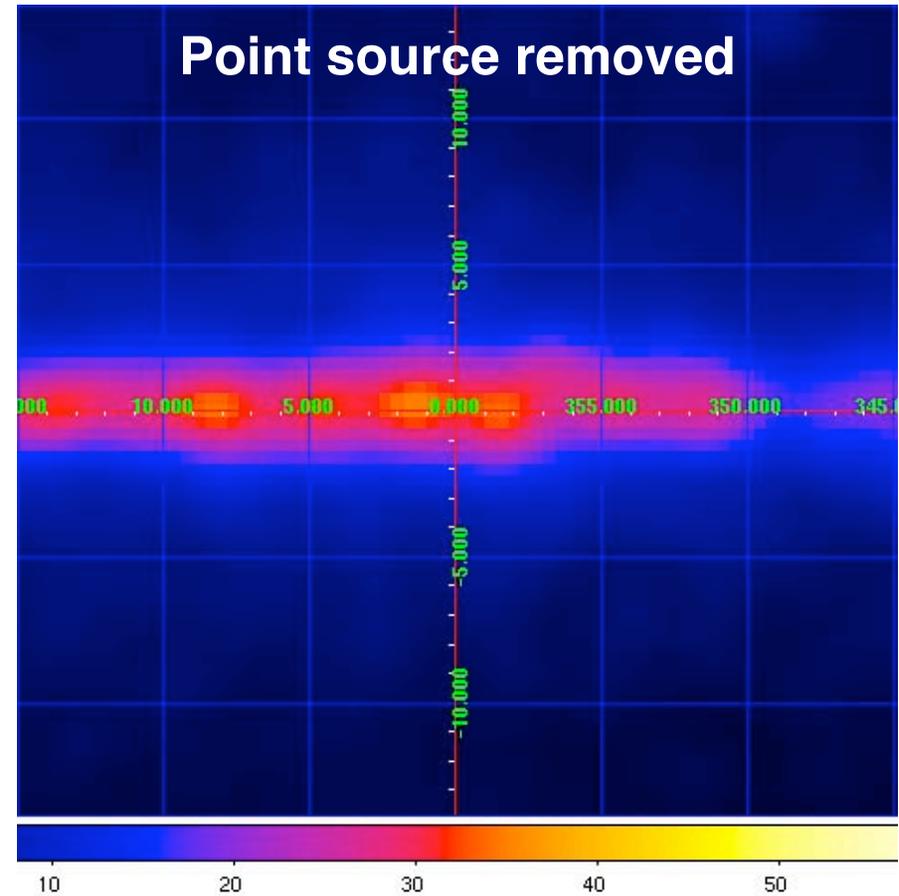
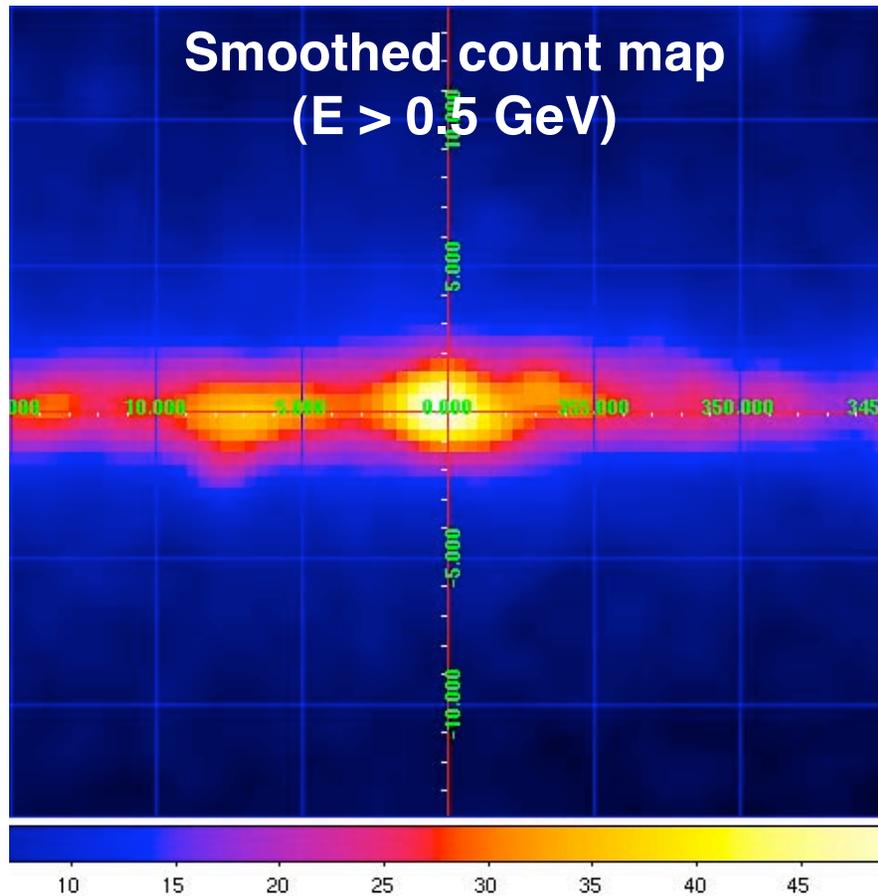


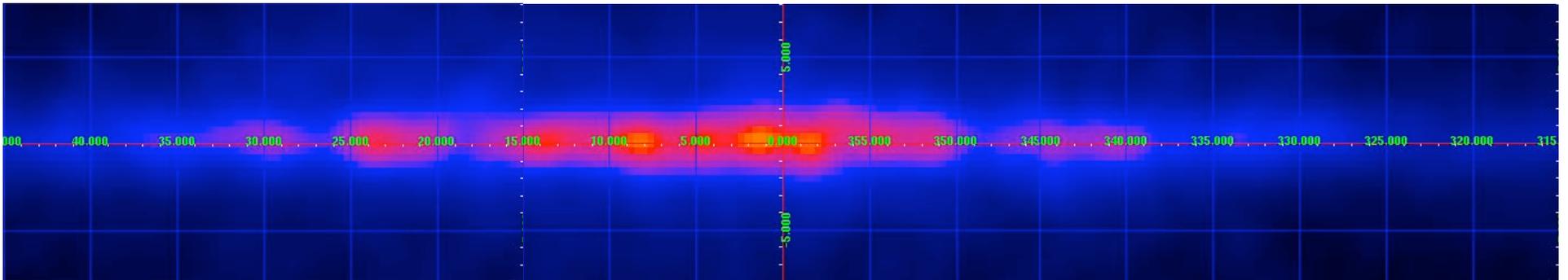
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Galactic Plane



- **3EG point sources are removed.**
 - » **$E > 0.5$ GeV**



- » **$0.5 > E > 0.2$ GeV**

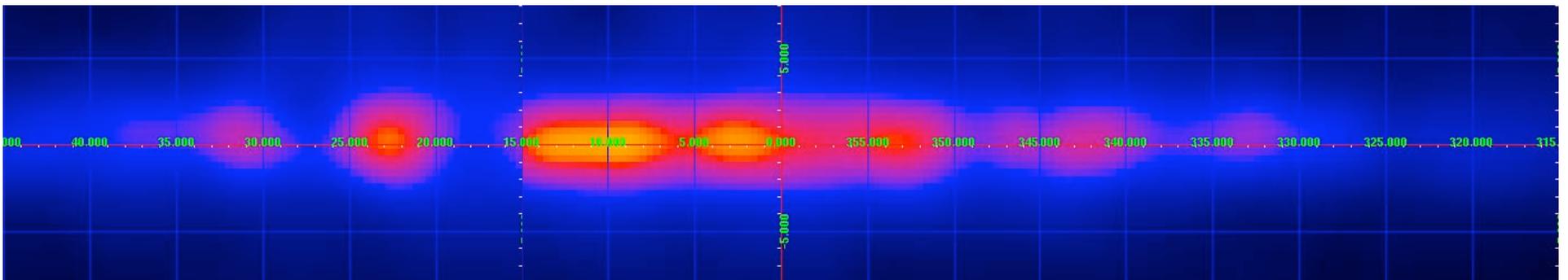


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Summary



- **Image deconvolution with event-by-event PSF is a promising tool study extended sources in EGRET and GLAST.**
- **Demonstration with EGRET sources.**
 - » **Improved image of LMC.**
 - » **Indication of an extended source or two point sources for 3EG J1234-1318.**
- **Further improvement and studies.**
 - » **Derive energy spectrum for each source.**
 - » **Understand systematic effects.**
 - **Statistical treatment of signal candidates.**